# Higgs-Exempt No-Scale Supersymmetry and its Experimental Signatures

David Morrissey

University of Michigan

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# Outline

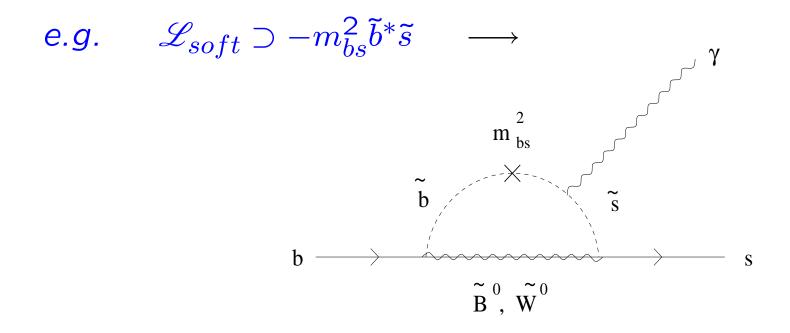
- The Supersymmetric Flavor Problem
- A Solution: Small Scalar Soft Terms
- RG Running and the Mass Spectrum
- Implications → Light Sleptons:
  - 1. Acceptable Bino dark matter
  - 2. Multi-lepton events at colliders.

#### Motivation: the SUSY Flavour Problem

- Low-scale supersymmetry (SUSY) is a well-motivated extension of the Standard Model.
- Supersymmetry can only be an approximate symmetry.

$$\mathscr{L} \supset \mathscr{L}_{soft} \supset -m^2 |\phi|^2 - A \phi^3 - M \lambda \lambda$$

• The  $m^2$  and A terms can induce too much flavour mixing.



### Approaches to the SUSY Flavor Problem

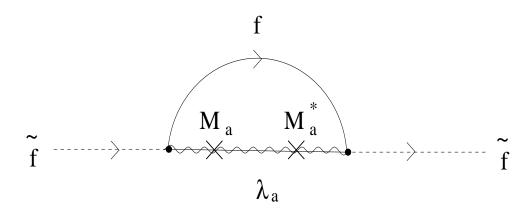
- 1. Introduce a new flavor symmetry that is broken above  $M_W$ . Such models are often very complicated.
- 2. Take  $\sqrt{|m^2|}$ ,  $A\gg 1000$  GeV, but  $M_{gaugino}\lesssim 1000$  GeV
  - → Split Supersymmetry

In doing so, we lose the explanation for  $M_W \ll M_{\rm Pl}$ .

- 3. Arrange for  $\sqrt{|m^2|}$ ,  $A \to 0$ , but  $M_{gaugino} \neq 0$ , at a scale much larger than  $M_W$ .
  - → Small Scalar Soft Terms (Splat Supersymmetry)
     This is the approach we will consider.

#### Small Scalar Soft Terms

- ullet Suppose all  $m^2$  and A soft terms vanish at some scale  $M_{input}\gg M_W$ , but the gaugino masses  $M_a$  do not.
- ullet Non-zero values for the  $m^2$  and A terms are generated as the theory is RG-evolved down to  $M_W$ .



 The soft scalar soft terms generated are nearly flavour universal, and do not generate too much flavour mixing.

# Obtaining Small Scalar Soft Terms

- These can arise in several ways:
  - 1. No-Scale models [Cremmer et al. '83, Ellis et al.'84]
    - A form of gravity mediation motivated by certain string theory compactifications.
  - 2. Gaugino mediated SUSY breaking [Chacko et al.'99, Kaplan et al.'99]
    - XD scenario with gauge multiplets in the bulk but chiral multiplets confined to branes.
  - 3. Strong conformal dynamics [Nelson+Strassler '00, Luty+Sundrum '01]
    - Interactions cause the scalar soft terms to flow to zero.

# Exempting the Higgs

• If all  $m^2$  and A terms vanish at  $M_{input}$ , the lightest superpartner is a charged slepton.

 $(M_{input} \leq M_{GUT} \simeq 2 \times 10^{16} \, {\rm GeV}$ , universal gaugino masses) [Schmaltz+Skiba '00, Komine+Yamaguchi '00, Baer *et al.* '02, Balazs+Dermisek '03]

- This can be a problem for cosmology.
- To avoid this, we consider an extension of the small scalar soft terms scenario that doesn't introduce flavour problems:

$$|m_{H_u}^2| \sim |m_{H_d}^2| \sim M_a^2 \gg m_{\tilde{f}}^2, \ A_f \quad \text{at} \ M_{input}.$$

 $\longrightarrow$  Higgs boson soft masses don't get squashed at  $M_{input}$  like the rest.

# Model Setup

We consider the following soft terms at scale

$$M_{input} = M_{GUT} \simeq 2 \times 10^{16} \text{GeV}$$
:

$$-m_{\tilde{f}}^2 = 0, \quad A_f = 0$$

$$-M_1 = M_2 = M_3 = M_{1/2} \leftrightarrow 1, 2, 3 = U(1), SU(2), SU(3)$$

$$-m_{H_u}^2$$
,  $m_{H_d}^2$  unfixed.

• The independent free parameters of the model are:

$$M_{1/2}, m_{H_u}^2, m_{H_d}^2, \tan \beta, sgn(\mu).$$

# Renormalization Group Evolution of Soft Terms

ullet For universal gaugino masses at  $M_{GUT}$  the electroweak scale gaugino masses are:

$$M_1 \simeq (0.41) M_{1/2},$$
 $M_2 \simeq (0.82) M_{1/2},$ 
 $M_3 \simeq (2.9) M_{1/2}.$ 

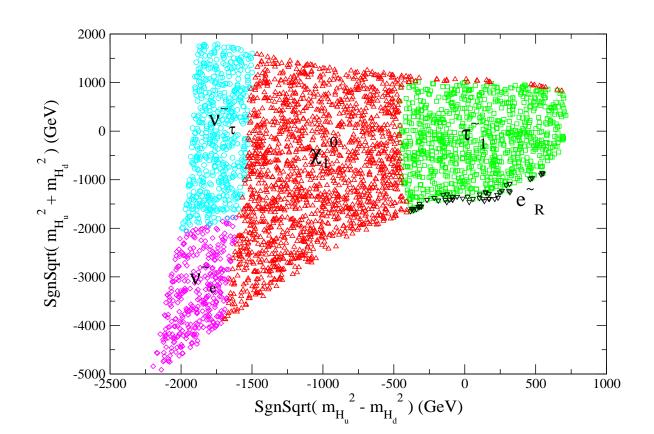
• The electroweak scale slepton soft masses are:

$$m_E^2 \simeq [(0.39)\,M_{1/2}]^2 - (0.055)\,S_{GUT}$$
 
$$m_L^2 \simeq [(0.64)\,M_{1/2}]^2 + \frac{1}{2}(0.055)\,S_{GUT}$$
 where  $S_{GUT} = (m_{H_u}^2 - m_{H_d}^2)_{GUT}$ .

• The low-scale squark masses are considerably larger.

# Lightest Superpartners

- $\tan \beta = 10$ ,  $M_{1/2} = 500$  GeV,  $sgn(\mu) > 0$ .
- Only for  $S_{GUT}=(m_{H_u}^2-m_{H_d}^2)<0$  is the LSP a neutralino.

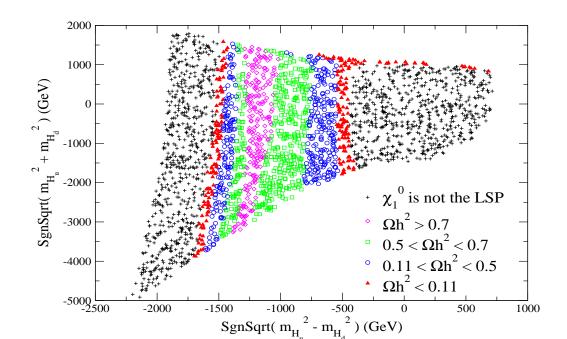


# Features of the Mass Spectrum

- ullet All masses scale with the input gaugino mass  $M_{1/2}$ .
- Lighter electroweak gauginos
  - ightarrow Higgs mass constraint requires  $M_{1/2} >$  300 GeV
- Lighter sleptons
  - → some sleptons are close in mass to the lightest gaugino
- Heavier squarks and gluino
  - → these help to push the Higgs mass up
- The light sleptons play a key role in the cosmology and collider signatures of the model.

# Light Sleptons #1: Neutralino Dark Matter

- With a universal gaugino mass  $M_{1/2}$ , the LSP is mostly  $U(1)_Y$  gaugino and tends to yield too much dark matter.
- If there is slepton close in mass to the (Bino-like) LSP, they can coannihilate.
- $\bullet$  tan  $\beta = 10$ ,  $M_{1/2} = 500$  GeV,  $sgn(\mu) > 0$

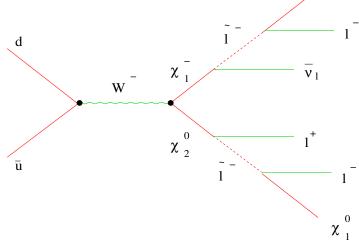


# Light Sleptons #2: Leptons at Colliders

• Light sleptons lead to leptonic events at colliders:

$$\chi^0_{2,3,4} o \tilde{\ell}^\pm \ell^\mp$$
 are kinematically allowed  $\chi^+_{1,2} o \tilde{\ell}^+ \, \nu, \, \tilde{\nu} \, \ell^+$ 

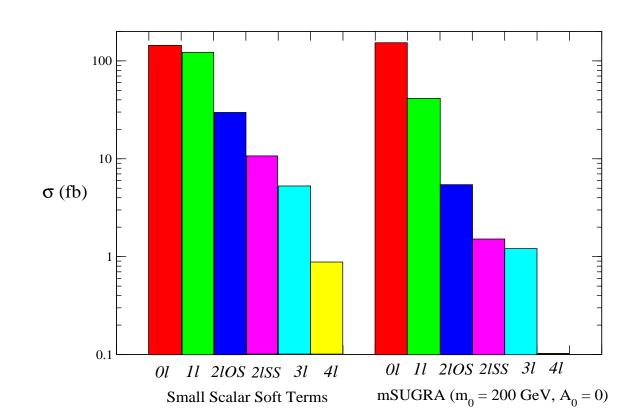
• e.g. Trileptons at the Tevatron



Using the set of cuts in [Baer et al '99] we find:

$$\sigma_{3\ell}\lesssim 0.5~{
m fb}$$
 (  $M_{1/2}=$  300 GeV,  $aneta=$  10, "HC2" cuts)  $\sigma_{bg}\simeq 0.5~{
m fb}$ 

- With small scalar soft terms, the LHC should be able to discover SUSY with 10  $fb^{-1}$  of data for  $M_{1/2}\lesssim 700$  GeV.
- Compared to other SUSY scenarios, the ratio of  $1\ell$  and multi- $\ell$  events to  $0\ell$  events is very large.
- For  $\tan \beta = 10$ ,  $M_{1/2} = 500 \, \text{GeV}$ , and simple cuts,



#### List of Cuts

- Events were simulated using ISAJET 7.74 [Baer et al. '06]
- "Lepton"  $\Rightarrow$  isolated e or  $\mu$  with  $p_T > 10 \, {\rm GeV}$ ,  $|\eta| < 2.5$ .
- ullet All Events:  $E_T > 200 \, {
  m GeV}$ ,  $S_T > 0.2$ ,  $n_{jets} \geq 2$ .
- $0 \ell$ :  $30^o < \Delta \phi(E_T, j) < 90^o$  with the nearest jet.
- 1  $\ell$ :  $p_T(\ell) >$  20 GeV,  $M_T(\ell, E_T) >$  100 GeV
- $\geq 2\ell$ :  $p_T(\ell_{1,2}) > 20 \text{ GeV}$ .

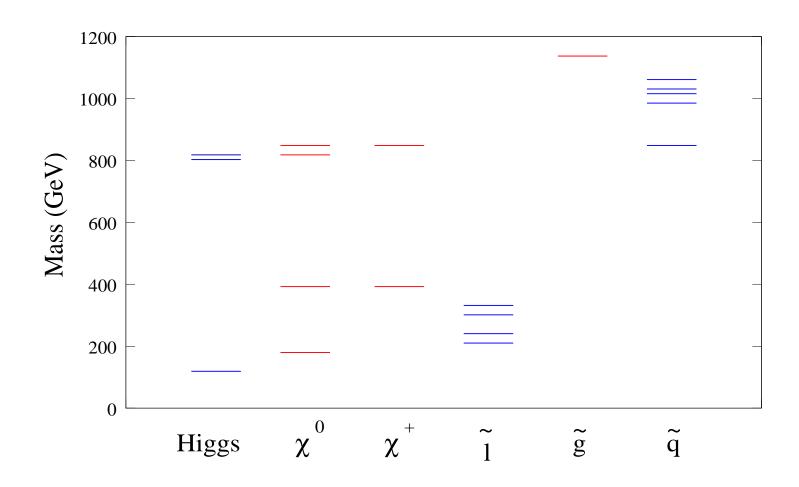
# Summary

- ullet Small scalar soft terms at the input scale  $M_{input}$  is a simple solution to the SUSY flavour problem.
- Allowing unsuppressed Higgs soft terms allows for a neutralino LSP and doesn't reintroduce a flavour problem.
- This scenario always yields very light sleptons.
  - ⇒ acceptable dark matter density through coannihilation
  - ⇒ many multi-lepton events at the LHC

# Extra Slides

# Sample Mass Spectrum

•  $\tan \beta = 10$ ,  $M_{1/2} = 500$  GeV,  $sgn(\mu) > 0$ .



# RG Evolution of Gaugino Masses

The RG evolution equation for the gaugino masses is

$$\frac{dM_a}{dt} \simeq -\frac{2b_a}{(4\pi)^2} g_a^2 M_a, \quad a = 1, 2, 3.$$

- $\bullet \Rightarrow M_a/g_a^2$  is approximately scale-independent.
- If  $g_3=g_2=g_1=g_{GUT}\simeq 0.72$  at  $M_{GUT}$ , the electroweak scale gaugino masses are

$$M_1 \simeq (0.41) M_{1/2},$$
 $M_2 \simeq (0.82) M_{1/2},$ 
 $M_3 \simeq (2.9) M_{1/2}.$ 

#### RG Evolution of Scalar Masses

The RG evolution equation for the soft scalar masses is

$$(4\pi)^2 \frac{dm_i^2}{dt} \simeq -8 \sum_a C_i^a g_a^2 |M_a|^2 + \frac{6}{5} g_1^2 Y_i S,$$

where

$$C_i^a = \begin{cases} 0, \frac{4}{3}, & a = 3 \\ 0, \frac{3}{4}, & a = 2 \\ \frac{3}{5}Y_i^2, & a = 1 \end{cases}$$

and

$$S=(m_{H_u}^2-m_{H_d}^2)+tr_F(m_Q^2-2m_U^2+m_E^2+m+m_D^2-m_L^2).$$

- ullet Because  $M_3$  and  $g_3$  grow large, the scalar quark masses also become large at the electroweak scale.
- The soft slepton masses are smaller,

$$m_E^2 \simeq [(0.39)\,M_{1/2}]^2 - (0.055)\,S_{GUT}$$
 
$$m_L^2 \simeq [(0.64)\,M_{1/2}]^2 + \frac{1}{2}(0.055)\,S_{GUT}$$
 where  $S_{GUT} = (m_{H_u}^2 - m_{H_d}^2)_{GUT}$ .

- $\bullet$  The mass of the lightest neutralino is usually set by  $M_1$ .
- The mass of the lightest slepton is usually less than  $min(\sqrt{m_L^2}, \sqrt{m_E^2})$ .

• If  $S_{GUT} \ge 0$  the lightest superpartner (LSP) is a mostly right-handed scalar lepton.

$$M_1 \simeq (0.41) M_{1/2}$$

$$\sqrt{m_E^2} \simeq \sqrt{[(0.39)M_{1/2}]^2 - (0.055) S_{GUT}}$$

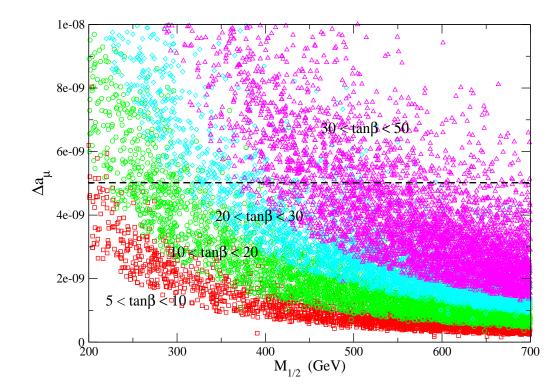
- This is problematic for cosmology.
- This is the difficulty for most no-scale type models.
- By allowing  $S_{GUT}=(m_{H_u}^2-m_{H_d}^2)<0$ , we can obtain a neutralino LSP.
- Higgs soft masses don't introduce a flavor problem.

# Light Sleptons #3: Muon Magnetic Moment

• The current value is [BNL-E821, PDG'06]

$$\Delta \left(\frac{g-2}{2}\right) = \Delta a_{\mu} = a_{\mu}^{exp} - a_{\mu}^{SM} = (2.2 \pm 1.0) \times 10^{-9}$$

• Light sleptons produce additional constributions to  $\Delta a_{\mu}$ .



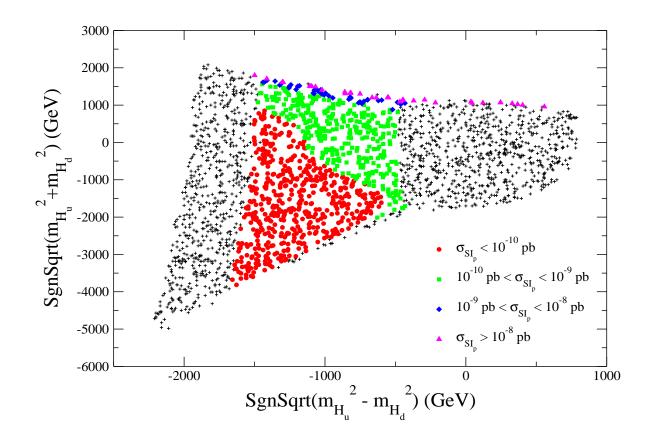
#### Direct Detection of Dark Matter

- The Earth encounters a flux of DM particles as it moves along with the galactic rotation.
- DM direct searches look for the interaction of dark matter particles with heavy nuclei.
- The limits from these searches are expressed in terms of an effective LSP-nucleon scattering cross section.
- Currently, the best limit is [CDMS II]

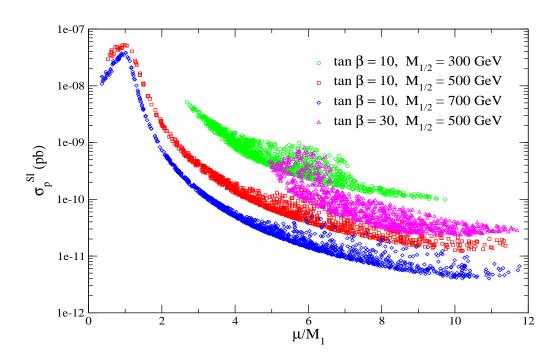
$$\sigma_p^{SI} < 10^{-6} - 10^{-7} \text{ pb},$$

for a DM particle of mass 50-1000 GeV.

 $\bullet$   $\tan\beta=$  10,  $M_{1/2}=$  500 GeV,  $sgn(\mu)>$  0  $(\sigma_p^{SI} \text{ was computed using DarkSUSY})$ 



- The direct detection rates lie below the CDMS II bound.
- ullet They increase as  $\mu/M_1$  grows smaller and the neutralino LSP develops a larger higgsino component.



ullet Upcoming direct detection experiments will probe down to about  $\sigma_p^{SI}\sim 10^{-9}$  pb.

#### Indirect Detection of Dark Matter

- Dark matter can also induce astrophysical signals.
- Neutralino LSPs can accumulate in the core of the Sun, where they can annihilate into neutrinos.
- Experiments such as Super-Kamiokande and AMANDA have placed bounds on this source of neutrino flux.
- ullet As for the direct detection bounds, the SSST signal is below the current bounds except when  $\mu/M_1$  is small.

# Signatures at the Tevatron

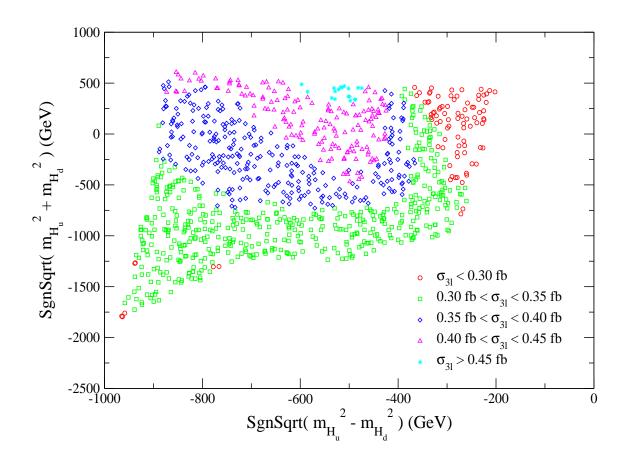
 The most promising channel at the Tevatron is the trilepton signal.

 $\begin{array}{c} \chi_{1}^{0} \\ \chi_{1}^{0} \\ \chi_{2}^{0} \\ \chi_{2}^{0} \\ \chi_{2}^{0} \end{array}$ 

• Since many of the sleptons are light, the branching ratio for  $\chi_2^0 \to \ell^- \tilde{\ell}^+$  is often large.

- Signal events were simulated using ISAJET 7.44 with the following set of cuts [Baer et al. '99]
  - 3 isolated leptons  $|\eta(\ell_{1,2,3})| < \text{2.5, } p_T(\ell_{1,2,3}) > \text{20, 15, 10 GeV}$
  - $-p_T > 25$  GeV.
  - lepton momenta not from  $W^\pm$  and  $Z^0$  decays  $12\,{\rm GeV} < M_{inv}(\ell^+\ell^-) < 81\,{\rm GeV} \ {\rm for}\ {\rm OS}\ {\rm SF}\ {\rm lepton}\ {\rm pairs}$  NOT  $60\,{\rm GeV} < m_T(\ell,\rlap/p_T) < 85\,{\rm GeV}$
- With these cuts, the main SM background comes from  $W^*Z^*$  and  $W^*\gamma^*$  production, and is about 0.5 fb.

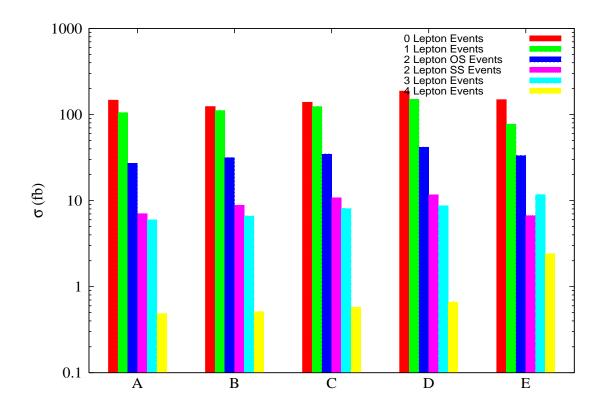
• For  $M_{1/2}=300$  GeV,  $\tan\beta=10$ , the signal is about 0.3-0.5 fb.



 $\Rightarrow$  possible hints from the Tevatron with  $10 fb^{-1}$  of data?

# Signatures at the LHC

- ullet Small scalar soft terms  $\to$  light sleptons  $\to$  leptonic events.
- The distinguishing feature of these scenarios is a high rate for multi-lepton events.
- $\bullet$  For tan  $\beta=10$ ,  $M_{1/2}=500\,\mathrm{GeV}$ , and simple cuts,



 $\bullet$  With small scalar soft terms, the LHC should be able to discover SUSY with 10  $fb^{-1}$  of data for  $M_{1/2}\lesssim$  700 GeV.

- Compared to other SUSY scenarios, the ratio of  $1\,\ell$  and multi- $\ell$  events to  $0\,\ell$  events is very large.
- The  $3\ell$  and  $4\ell$  channels are particularly distinctive. Signals:

$$\sigma_{3\ell} = 5 - 10 \, fb$$
  $\sigma_{4\ell} \gtrsim 0.5 \, fb$ 

#### SM Backgrounds:

$$\sigma_{3\ell}^{bg} = 0.1 \, fb$$
  $\sigma_{4\ell}^{bg} \simeq 0.002 \, fb$